# **Physics**

# **Todays Agenda**

- Reference frames and relative motion.
- Uniform Circular Motion



#### **Inertial Reference Frames:**

- A **Reference Frame** is the place you measure from.
	- I It's where you nail down your (x,y,z) axes!
- An Inertial Reference Frame (IRF) is one that is not accelerating with respect to the " fixed stars".
	- **We will consider only IRF's in this course.**
	- The earth is a pretty good IRF even though it is accelerating (rotating) w.r.t. the stars slightly.
- Valid IRF's can have fixed velocities w.r.t each-other.
	- More about this later when we discuss forces.
	- **For now, just remember that we can make** measurements from different vantage points.

#### **Relative Motion**

● Consider a problem with **two** distinct IRF's: ❵ **An airplane flying on a windy day.**

A pilot wants to fly from Austin to Dallas. Having asked a friendly student, she knows that Dallas is 120 miles due north of Austin. She takes off from the Airport at noon. Her plane has a compass and an air-speed indicator to help her navigate.

- **The compass allows her to keep the nose of the plane** pointing north.
- **The air-speed indicator tells her that she is traveling at Fig. 2.1.** 120 miles per hour **with respect to the air**.

## **Relative Motion...**

- The plane is moving north in the IRF attachd to the air:
	- **i**  $V_{p,q}$  is the velovity of the plane w.r.t the air.



#### **Relative Motion...**

- But suppose the air is moving east in the IRF attached to the ground.
	- **I**  $V_{a,g}$  is the velocity of the air w.r.t the ground (i.e. wind).



# **Relative Motion...** ● What is the velocity of the plane in an IRF attached to the ground? **i**  $V_{p, g}$  is the velocity of the plane w.r.t the ground.

**Vp, g**

## **Relative Motion...**

$$
V_{p,g} = V_{p,a} + V_{a,g}
$$

**Is a vector equation relating the airplanes** velocity in different reference frames.



# **Uniform Circular Motion**



- What does it mean?
- How do we describe it?
- What can we learn about it ?



#### **How can we describe UCM?**

- In general, one co-ordinate system is as good as any other:
	- **B** Cartesian:
		- » (x,y) [position]
		- $\rightarrow$   $(V_x, V_y)$  [velocity]
	- ❵ Polar:
		- » (R,θ) [position]
		- »  $(V_R, \omega)$  [velocity]
- In UCM:
	- **B** R is constant (hence  $V_R = 0$ ).
	- ❵ ω (anguar velocity) is constant.
	- ❵ Polar co-ordinates are a natural way to describe UCM!



#### **Polar Coordinates:**

- The arc length **s** (distance along the circumference) is related to the angle in a simple way:
	- $s = R\theta$ , where  $\theta$  is the angular displacement.
	- $\mathbf{l}$  units of  $\theta$  are called radians.

● For one complete revolution:  $2\pi R = R\theta_c$ 

 $\overline{\theta}_{c} = 2\pi$ θ has period 2π.

**R**  $\setminus$  . **v**  $\mathbf{y}_1$  and  $\mathbf{y}_2$  $(x,y)$ **s** θ

 $\frac{1}{2}$  1 revolution =  $2\pi$  radians.

SDW - Physics - VHS

 $\mathbf{x}$  and  $\mathbf{x}$ 



## **Polar Coordinates...**

 $\bullet$  In cartesian co-ordinates we say velocity  $dx/dt = v$ .

 $\}$   $x = vt$ 

• In polar coordinates, angular velocity  $d \theta/dt = \omega$ .  $\theta = \omega t$ 





● Even though the speed is constant, velocity is **not** constant since the direction is changing.

❵ Consider average acceleration in time ∆t ∆**v** / ∆t



- Even though the speed is constant, velocity is **not** constant since the direction is changing.
	- ❵ Consider average acceleration in time ∆t ∆**v** / ∆t





 $\overbrace{w\Delta t}$  Move ∆**v** to average time of  $\overbrace{w\Delta t}$  the interval  $\Delta t$ .

● Even though the speed is constant, velocity is **not** constant since the direction is changing.

❵ Consider average acceleration in time ∆t ∆**v** / ∆t



seems like ∆**v** (hence ∆**v**/∆t ) points toward the origin !

● Even though the speed is constant, velocity is **not** constant since the direction is changing.

**E** As we shrink  $\Delta t$ ,  $\Delta v / \Delta t$  dv/dt = a



#### **Aside: Polar Unit Vectors**

● We are familiar with the cartesian unit vetors: **i j k**



## **Centripetal Acceleration**

• Must be accelerating if direction is changing: ❵ Centripetal Acceleration !!



 $\frac{\Delta V}{V} = \frac{\Delta R}{R}$  $\overline{\mathsf{R}}$  $\Delta R$  $R$  and  $\parallel$  and  $\parallel$ Similar triangles:  $\frac{\Delta V}{\Delta}$ 

But  $\Delta R = v \Delta t$  for small  $\Delta t$ 

Similar triangles: 
$$
\frac{\Delta v}{V} = \frac{\Delta V}{R}
$$
  
But  $\Delta R = v\Delta t$  for small  $\Delta t$   
So:  $\frac{\Delta V}{V} = \frac{v\Delta t}{R}$   $\frac{\Delta V}{\Delta t} = \frac{v^2}{R}$ 

$$
a = \frac{v^2}{R}
$$

## **Centripetal Acceleration**

- UCM results in acceleration:
	-
	-

**i** Magnitude:  $a = v^2 / R$   $(= \omega^2 R \text{ since } v = R\omega)$ **B** Direction:  $\hat{r}$  (toward center of circle)



#### **Example: Propeller Tip**

• The propeller on a stunt plane spins with frequency  $f = 3500$  rpm. The length of each propeller blade is  $L=$ 80cm. What centripetal acceleration does a point at the tip of a propeller blade feel?



#### **Example:**

● First calculate the angular velocity of the propeller:

1 
$$
\text{rpm} = 1 \frac{\text{rot}}{\text{min}} \times \frac{1}{60} \frac{\text{min}}{\text{s}} \times 2 \pi \frac{\text{rad}}{\text{rot}} = 0.105 \frac{\text{rad}}{\text{s}} = 0.105 \text{ s}^{-1}
$$

 $\frac{1}{2}$  so 3500 rpm means  $\omega = 367 \text{ s}^{-1}$ 

• Now calculate the acceleration.  $a = \omega^2 R = (367s^{-1})^2 \times (0.8m) = 1.1 \times 10^5 m/s^2$  $= 11,000$  g

❵ direction of **a** is toward propeller hub (-**r** ). **^**

# **Example: Acceleration at Equator.**

● What is the centripetal acceleration experienced by a person standing on the earth's equator, due to the earth's rotation.



• Recall that the radius of the earth is  $R_e = 6.35 \times 10^6$  m.



#### **Example: Newton & the Moon**

- What is the acceleration of the Moon due to its motion around the earth?
- What we know (Newton knew this also):
	- $\overline{I} = 27.3$  days = 2.36 x 10<sup>6</sup> s
	- $R = 3.84 \times 10^8$  m

 $R_F = 6.35 \times 10^6 \text{ m}$ 

(period  $\sim$  1 month) (distance to moon) (radius of earth)



#### **Moon...**

● Calculate angular frequency:

**WIO**  
\nalculate angular frequency:  
\n
$$
\frac{1}{27.3} \frac{rot}{day} \times \frac{1}{86400} \frac{day}{s} \times 2\pi \frac{rad}{rot} = 2.66 \times 10^{-6} \text{ s}^{-1}
$$

 $\bullet$  So  $\omega = 2.66 \times 10^{-6}$  s<sup>-1</sup>.

● Now calculate the acceleration.

 $a = \omega^2 R = 0.00272 \text{ m/s}^2 = .000278 \text{ g}$ 

*l* direction of **a** is toward center of earth (-**r**<sup> $\uparrow$ </sup>).

#### **Moon...**

- So we find that  $a_{moon}$  /  $g = .000278$
- Newton noticed that  $R_E^2/R^2 = .000273$



- This inspired him to propose the **Universal Law of Gravitation:**  $F_{Mm} = GMm / R^2$
- What if our solar system was more complicated...
	- ❵ Would early scientists have been as successful??
	- ❵ Would science have evolved differently??

## **Recap for today:**

- Reference frames and relative motion.
- Uniform Circular Motion
- Look at Textbook problems